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1	SHOCKED GAS TEMPERATURE MEASUREMENTS BY INFRARED MONOCHROMATIC RADIATION
2	PYROMETRY N66-18342 (THRU)
3	by M. R. Lauver, J. H. Hall, and F. E. Belles (PAGES) (CODE) 33
4	ABSTRACT (CATEGORY
5	18342 The temperature history of a carbon dioxide-argon mixture was followed
, 6	by an infrared momochromatic radiation pyrometer technique after incident
7	and reflected shocks. Temperatures as high as 3600° K were measured.
8	One-dimensional wave theory calculations of the incident and reflected
9	temperatures were found to be in good agreement with the experimental
10	results. Detailed time-temperature studies showed the effect of attenuation
11	and dissociation. A MARIE TO MACA OFFICE MACA PROPERTY
12	INTRODUCTION
13	The shock tube has found one of its prime applications in the study
14	of high-temperature rate processes in gases. Commonly, these processes
15	have a strong temperature dependence, and it is therefore necessary to
16	know the temperature of the experiment quite accurately. This need can be

met nicely so long as incident shock waves are used to heat the gas,

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because one-dimensional theory, plus the thermodynamic properties of the gas, can be used to calculate the shocked-gas temperature from the measured velocity of the wave. Very often, however, it is convenient or even necessary to study processes behind reflected shocks, taking advantage of the higher temperatures and pressures which can be produced without unduly straining the capacities of conventional shock tubes. In that event, it is no longer so clear that the temperature can be calculated accurately; interactions of the reflected shock with the boundary layer (Ref. 1) and with the pressure gradient created by attenuation of the incident shock (ref. 2) introduce ambiguities.

This problem has been widely recognized, and measurements of reflected-gas temperature have recently begun to appear in the literature. For the most part they were made by the line-reversal method (ref. 3,4), although in one case the rate of a chemical reaction was used as an indication of temperature (ref. 5). The purpose of the present work was to apply still another technique. We have determined the temperature from simultaneous measurements of infrared spectral emission and emissivity.

1	Among the advantages of the infrared method are the following: (1) There
2	is no need to introduce sodium salts or other sources of emission; (2) good
3	time resolution is possible; and (3) a very wide temperature range can
4	be covered, not limited by the available brightness temperature of a
5	light source, as in the reversal method.
₋ 6	This paper reports temperatures measured behind reflected shocks in
7	a gas that is typical of mixtures likely to be used in chemical studies:
8	10% CO ₂ -90% Ar. The temperature range behind reflected shocks was
9	2000° to 3600° K, although temperatures as low as 1000° K were measured
10	behind incident waves. Temperatures determined at short times (50 to 100 μ
11	sec) after the reflected wave passed the observation station are compared
12	with values calculated from the velocity of the incident wave. Two detailed
13	temperature-time records are also presented; these show the effects of
14	shock attenuation on both the incident- and the reflected-shock temperature.
15	APPARATUS AND EXPERIMENTAL PROCEDURE

The infrared monochromatic radiation (1MRA) method of gas temperature

Basis of Emission-Absorption Pyrometry

measurement has been described elsewhere (ref. 6, 7). It is based on the fact that the intensity of radiation emitted by a hot gas depends on the gas temperature and on the number of gaseous entities (molecules) free radicals, or atoms) emitting radiation. Consequently, by measuring the absolute intensity of radiation (spectral emission) and the relative number density of emitters (spectral emissivity), the temperature of a gas can be calculated. Figure 1 shows schematically the information recorded in a shock-tube experiment, and the way in which it is reduced to a temperature measurement.

The measurement of spectral emission is straightforward. The infrared detection system response to the hot gas emission $(V_g, f_ig. 1)$ is compared with its response to a calibrated standard infrared source (V_g) , and converted to absolute units by means of calibration factors. The result is a spectral emission value (I_g) in terms of watts $cm^{-2}\mu^{-1}$ ster⁻¹ at a specific wavelength. The effect of the number of emitters is accounted for by absorption spectroscopy at the same wavelength. The reduction in signal strength due to transmitted light, from V_0 to V volts, determines

the spectral absorptivity $(1-\frac{V}{V_0})$, which is equal to the spectral emissivity, e.

According to Kirchoff's radiation law, the ratio of the spectral emission of the hot gas to its spectral emissivity is numerically equal to the radiation (I_b) that would be emitted by an ideal blackbody radiator at the same temperature as the gas:

$$\frac{I_{g}}{e} = I_{b} \tag{1}$$

Planck's law relating temperature to blackbody spectral emission is

$$I_b = c_1 \lambda^{-5} / (e^{\frac{C_2}{\lambda} T} - 1)$$
 (2)

Thus, inasmuch as C_1 and C_2 are physical constants and I_b is measured at the wavelength λ , the temperature is determined.

Apparatus for Emission-Absorption Pyrometry

A schematic diagram of the shock tube and associated equipment is given in figure 2. A constant intensity of infrared radiation from a glower source is mechanically chopped at 80 kilocycles per second and sent through the spephire windows of the shock tube into a prism monochromator and a liquid-nitrogen-cooled indium antimonide detector. The shocked gases emit radiation which passes out of the shock tube into the same radiation

detection system. Its output is amplified by a directly-coupled circuit

providing a measure of the gaseous emission intensity (the glower intensity

is negligible in comparison to that of emission). The glower signal

intensity is determined with a capacitor-coupled, tuned circuit of high

gain. With these amplifiers, the detection system distinguishes between

the emitted and transmitted light, although both are presented simultaneously

to the one detector, and sends the resulting signals to different oscilloscope

beams for recording.

Calibration of Phyrometer:

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The IMRA apparatus was calibrated by a determining the transmittance of the shock tube window and the absolute radiance of the internal secondary-standard source (B in Fig. 2), which was a tungsten-ribbon lamp.

The window was found to pass 82 percent of 4.5 μ radiation falling on it. This was the wavelength used for the temperature measurements. Although the center of the asymptotic stretching band of carbon dioxide is at 4.3 μ , the longer wavelength is much more desirable. At 4.5 μ ,

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- only the hot gas in the shock tube is optically active, while the cool
- 2 carbon dioxide in the room and in the shock-tube boundary layer is almost
- 3 completely transparent.
- 4 The internal standard was compared with a source which had been
- 5 calibrated at the Bureau of Standards. It was found to emit 2.47 watts
- $_{1}$ 6 cm $^{-2}\mu^{-1}$ ster $^{-1}$ at 2.2 μ . The relative sensitivity of the optical system
- 7 and detector at 4.5 and 2.2 μ was then determined by viewing a blackbody
- 8 source of known temperature at the two wavelengths.
- 9 With the foregoing calibration factors at hand, the absolute spectral
- 10 emission, I_a , of the hot gas at 4.5 μ is readily determined from the
- 11 following formula:

12
$$I_g = (V_g/V_s) (1/R)(1/t_w) (I_s)$$
 (3)

- 13 where
- V_{q} = amplitude of emission signal, volts
- $V_s = amplitude$ of signal from internal standard source, volts
- 16 R = sensitivity of the system at 4.5 μ relative to that at 2.2 μ
- $t_w = \text{transmittance of shock-tube window at } 4.5 \mu$.

 $I_s = intensity of internal standard at 2.2 <math>\mu$.

This formula is valid provided the amplification and the slit width are the same when both $\,^{\rm V}_{\rm g}\,$ and $\,^{\rm V}_{\rm s}\,$ are recorded. If they are not the same, the appropriate correction terms must be applied.

Shock Tube

The shock tube was of rectangular cross section, 37mmx74mm. Two circular sapphire windows, 28mm in diameter, were flush-mounted opposite one another across the longer dimension. The midpoint of the windows was 3.94 meters from the polyester plastic diaphragm which separated this driven section of the shock tube from the driver section. The midpoint of the windows was 179 mm from the downstream end wall of the driven section. This distance to the end wall could be reduced to 27mm or to 78mm by means of close-fitting plugs. The driver gas was helium. The diaphragms were pressure-burst.

Timing, Pressure, and Recording Instrumentation

Four thein-film resistance gauges were mounted upstream of the windows and one at the same axial position as the centerline of the windows.

1 These gauges marked the trime of incident (and in some cases, reflected) shock arrival at the five positions, relative to the first, or trigger, 2 3 position. The signals were displayed on one beam of a dual beam The other beam displayed the output of a quartz pressure oscilloscope. 4 This transducer was also located at a position corresponding 5 transducer. to the centerline of the windows. A second dual beam oscilloscope was 6 triggered to display absorption and emission levels from the gas as 7 measured by the IMRA apparatus. Timing pulses from a crystal-controlled 8 9 secondary frequency standard were recorded on all four oscilloscope beams 10 for each run. A calibration signal for the emission level interrupted 11 about 1100 times per second for identification and a calibration signal 12 for the pressure were also recorded each time. Figure 3 shows the oscillograph 13 records of infrared absorption and emission for a single experiment. 14 important features of this typical record are: (1) the large changes of 15 intensity of transmitted and emitted Light upon passage of the shock waves, 16 (2) the relatively noise-free signals; and (3) the quick recovery of the 17 tuned amplifier in the absorption channel from the ringing induced by

L	step-changes,	permitting	meaningful	readings	to	be	made	äbout	50	μ sec
>	after nassage	of the sho	ck							

It will be noted in figure 3 that the gas behind the reflected shock was practically black in this particular run. That is, it absorbed all of the radiation from the glower source. In many runs of this sort, the portion of the absorption trace after the reflected shock was simultaneously displayed on another oscilloscope at much higher gain, so that the absorptivity could be accurately determined.

Test Gas

A commercially prepared argon-carbon dioxide mixture was used without further treatment. It analyzed 9.9% carbon dioxide by volume, the balance, argon. A few tests were made with this gas diluted to 6.7% carbon dioxide in argon and with a commercially prepared argon-carbon dioxide mixture which analyzed at 1.1% carbon dioxide, the balance, argon.

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RESULTS AND DISCUSSION

The results of determinations of gas temperatures behind incident shocks are presented in fig. 4, where they are compared with temperatures calculated by simple shock theory from measured shock velocities. The IMRA temperatures were actually measured at several 10-µsec intervals for each experiment, starting at about 50 µsec laboratory time when the signals first became readable. Each point in figure 4 was obtained by extrapolating such data back to time zero, the instant when the shock passed the center of the windows. This was done to eliminate effects of attenuation on the temperature.

The agreement is good, and as expected, does not seem to be affected by the ${\rm CO}_2$ concentration over the range studied.

Despite this general agreement between measured and calculated incident temperatures, there is nevertheless a good deal of scatter evident in figure 4. Part of this scatter is traceable to contamination of the shock-tube windows, but most of it is due to a drifting type of instability of the infrared detector. From run to run, this instability produced

1	uncertanities	as	large as	10	percent	in	the	first	term	οf	equation	(3)	١.
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- 2 the result is an uncertainty of only about 30° at temperatures near 1000° K,
- but this grows to nearly 300° at temperatures near 4000° K.

Inasmuch as the main purpose of the work was to measure reflected-shock temperatures at levels above 2000° K, it was desirable to reduce the expected scatter in the results as much as possible. This was done by using the incident shock wave as an internal standard for each run. The temperature calculated from the measured shock velocity at the window position, and the measured emissivity of the gas behind the incident wave, were both assumed to be correct. The value of spectral emission I_g required to satisfy equations (1) and (2) was then calculated. This value, and the measured voltages V_g (incident shock) and V_s , were inserted in equation (3) and used to calculate a lumped instrument factor, $(I/R \times I/I_w)I_s$.

In each run, then, this individually-determined factor was used to convert measured voltages into a series of reflected-shock temperatures, starting at about 50 μ sec behind the shock and determined at approximately 10- μ sec intervals. These data were extrapolated to time zero, the instant

the reflected shock passed the center of the windows. The resulting

temperatures are plotted as ordinates in figure 5; as abscissas, we have

used the reflected temperatures calculated from the velocity of the incident

shock as it passed the windows, using Markstein's method (ref. 8).

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This means of obtaining calculated reflected-shock temperatures for comparison with measured values is clearly rather arbitrary. Its most obvious shortcoming is that it does not allow for the attenuation of the incident wave. However, it is important to see how well the temperatures calculated in this simple way agree with the measured values. Figure 5 shows that they agree surprising well. The points obtained with the end wall at its most remote position, 179mm from the windows, tend to be high, and those obtained with the end wall 78mm away tend to be low. This behavior is not inconsistent with that noted in reference 9, which reported that the reflected-shock pressure behaved as if the shock detelerated and then accelerated as it receded from the end wall. in a grosser sense, the agreement between measured and simply-calculated temperatures is very grafifying. Of the 27 measurements shown in figure 5, 1 14 are within 100° of the line, and only 4 miss it by more than 200° K.
2 We do however, believe that deviations lagger than 100° are real, and
3 that they reflect actual departures of the shocks from ideal one-dimensional behavior.

An example of the detail which it is possible to obtain by the IMRA method, is given in f igure 6, where the temperature history of a shock is followed at 10 to 20 μ sec intervals for 700 μ sec after its passage.

The rise in temperature behind the incident shock is accounted for by attenuation effects. This shock was determined to be attenuating in velocity at the rate of 1.5×10^{-4} mm/ μ s/mm $\left(=\frac{dU}{dg}\right)$, and therefore, for 10^{-6} sec, $\frac{dU}{U}=1.5 \times 10^{-4}$. Strong shock theory predicts $\frac{dT}{T}=\frac{dP}{P}=2\frac{dU}{U}$. (Calculations at temperatures of the order of 1300° K, show $\frac{dT}{T}\cong\frac{7}{4}\frac{dU}{U}$ and $\frac{dP}{P}\cong2\frac{dU}{U}$). Pressure measurements gave $dP/P=3.8 \times 10^{-4}$ behind the incident shock, somewhat higher than would have been indicated by attenuation. The IMRA results in f igure 6 show $dT/T=3.0 \times 10^{-4}$, in good accordance with the attenuation prediction.

Other observers (ref. 2) have noted a rise in pressure behind reflected

- shocks. This pressure rise, converted into an isentropic temperature
- 2 change by the equation $T = \left(\frac{P}{P_o}\right) + \frac{Y-1}{F} T_o$, corresponds very well with
- 3 the infrared pyrometer results of figure 6.
- 4 A second example of detailed analysis is presented in **f**igure 7.
- 5 Here the temperature following the passage of the reflected shock tends
- 6 to rise, due to isentropic compression, and tends to decrease due to
- 7 dissociation according to the equation
- 8 $co_2 \rightarrow co + \frac{1}{2} o_2$, $\Delta H = 65.8 \text{ K cal}$
- 9 (The rate of oxygen atom recombination is sufficiently high so that the
- concentration of 0_2 is much greater than that of 0 at this temperature).
- Using: (1) 2.9×10^7 cc moles $^{-1}$ sec $^{-1}$ °K $^{-1}$ for the dissociation rate of
- 12 CO, at 3220°K (ref. 110) and for the subsequent temperatures after the
- reflected shock wave passage, (2.) the isentropic temperature corrections
- from the measured pressures, and (3.) negligible back reaction of the CO
- and 0 to form ${\tt CO}_2$, theoretical temperature calculations were made. As
- with the simpler case of f igure 6, the agreement of theory and results is
- gratifying. In this case, figure 7, the attenuation comparisons behind

- 1 the incident shock are $\frac{dU}{U} = 2.6 \times 10^{-4}$, $\frac{dP}{P} = 4.3 \times 10^{-4}$, and $\frac{dT}{T} = 3.4 \times 10^{-4}$.
- 2 CONCLUSIONS

- 1. Infrared pyrometry of a carbon dioxide-argon gas mixture yielded

 4 temperatures in agreement with those calculated from the incident shock

 5 speed by common shock theory, from 1100 to 3600° K.
 - 2. A comparison of reflected shock temperatures measured 27, 78, and

 179mm after reflection with temperatures calculated from one-dimensional

 shock theory shows latter are generally good but sporadically may be in error by large amounts.
 - 3. Attenuation of the speed of an incident shock wave was accompanied by changes in the gas pressure and temperature in the period following the passage of the shock. The relationships between these changes were in general agreement with strong shock theory.
 - 4. The change in gas temperature with time after a reflected shock was adequately calculated from the dissociation rate of carbon dioxide and the pressure history of the gas mixture.
- 5. These data were preliminary in nature. It is expected thatrefinements in apparatus and technique will yield much improved data.

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1	R F	P F K	FNCES

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FIG I BALLS OF IMPA

TEMPERATURE MENCOLEMENT

Betwee shork - after shork -

Emissivity, e= Absorptivity = 1-(V./Vo)

Kirchoffs Radiation Law:
Spectral Emission Le : partial Emission of
Spectral Emissionity e : Blackbody at Same

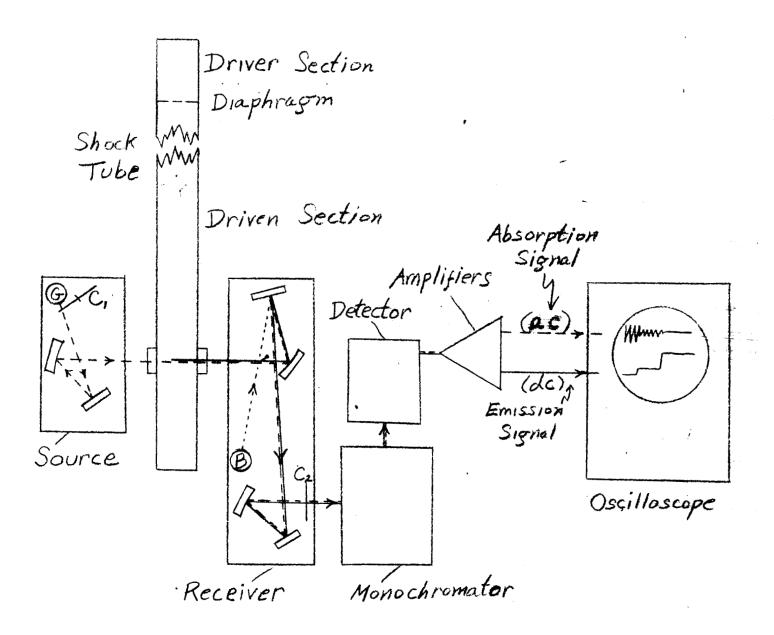
I, = c, 2-5/(c (2/2T-1)) 3. Planck's Low:

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Figure 2. - Schematic Diagram of a Monochromatic Radiation Pyrometer.



G Glower

B Internal
Standard

C, 80 KCps Chopper

C2 1.1 Kcps Chopper

FIG.3 Typical Oscilogram

BEFORE

TIKIDENT

SHOCK

AFTER AFTER REFLECTED

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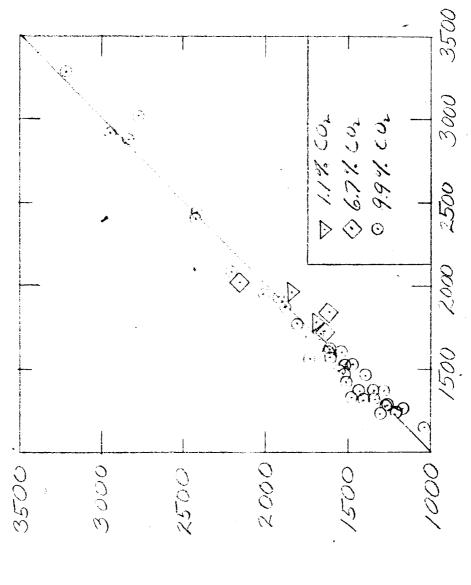
SHOCK

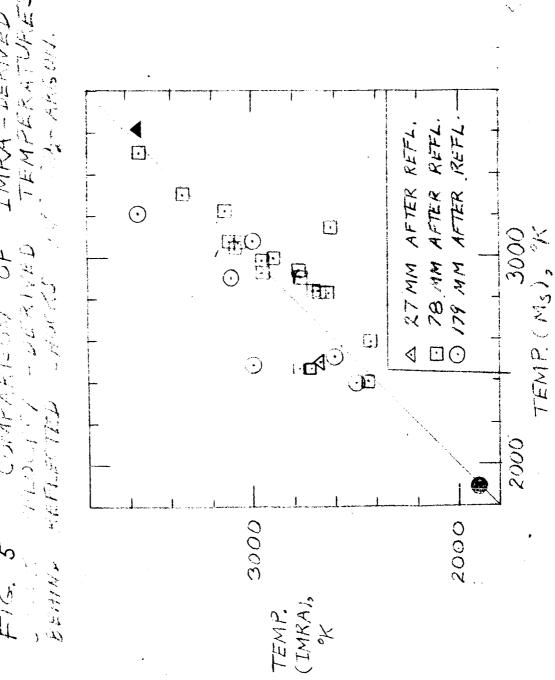
ABSORPTION

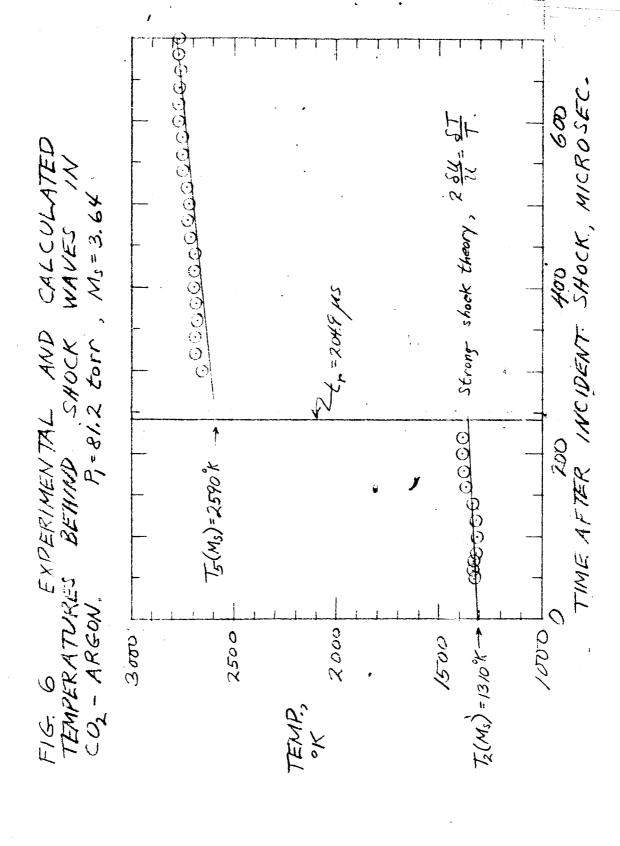
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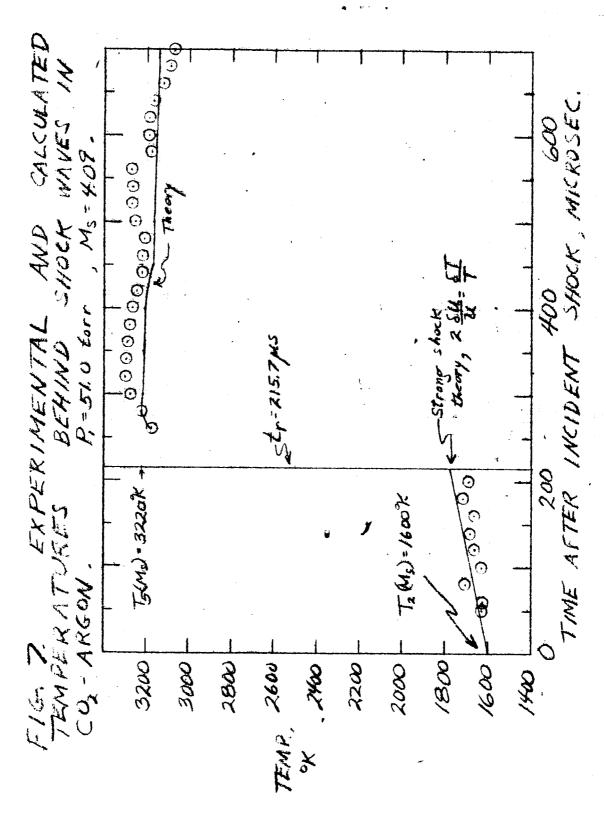
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TEMPERATURES BEHIND CO4-48608 FIG. 4 COMPAKISON C SHOCK VELOCITY - SERIVED TNCIDENT SHOCKS IN









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